



ELSEVIER

30 October 2000

Physics Letters A 276 (2000) 1–7

---

---

PHYSICS LETTERS A

---

---

www.elsevier.nl/locate/pla

# The speed of quantum information and the preferred frame: analysis of experimental data

Valerio Scarani \*, Wolfgang Tittel, Hugo Zbinden, Nicolas Gisin

*Group of Applied Physics, University of Geneva, 20, rue de l'Ecole-de-Médecine, CH-1211 Geneva, Switzerland*

Received 10 August 2000; accepted 18 September 2000

Communicated by P.R. Holland

---

## Abstract

The results of EPR experiments performed in Geneva are analyzed in the frame of the cosmic microwave background radiation, generally considered as a good candidate for playing the role of preferred frame. We set a lower bound for the speed of quantum information in this frame at  $2 \times 10^4 c$ . © 2000 Elsevier Science B.V. All rights reserved.

*Keywords:* Optical EPR experiments; Preferred frame; Cosmic background radiation

---

## 1. Introduction

The tension between quantum mechanics (QM) and relativity manifests itself in two classes of theoretical problems. The first class of problems can be labelled “the search for a covariant description of the measurement process”. Possibly the best-known example is the impossibility of a causal description of the collapse in an EPR experiment that would be valid in all frames; but there are many other examples, even for one-particle measurements, as widely discussed by Aharonov and Albert [1,2]. The second class of problems is linked with some structural problems of quantum relativistic theories, like the definition of a position operator that fulfills “basic” requirements [3–5].

In this paper, we are concerned with the first of these classes. Actually, from an “orthodox” standpoint, these problems have been solved [1,6,7]: two

observers, each using quantum prescriptions, predict the same final probabilities — recall that in its orthodox interpretation, QM deals only with probabilities, while the state vector and its evolution are not endowed with reality [8]. But several physicists are not satisfied with this solution, for different reasons [9]. For us, the tension between the notion of event that appears in relativity, and the reversible evolution of the quantum state, may be a guide for new physics.

The introduction of a preferred frame (PF) is a way out of the first class of problems, that would allow a realistic (obviously non-local) description of the quantum measurement [10]. Moreover, it seems that the PF would also be a way out of the second class of problems [3,4]. To our knowledge, the introduction of a PF is still an intellectual tool (or trick): no experiments are planned or even proposed that aim to falsify this hypothesis [11]. However, an experiment can be easily conceived that falsifies a *joint hypothesis*: suppose that there is a PF *and* that in the PF the speed of quantum information [12] is finite, though superluminal [13]. Then, if quasi-simultaneity in the

---

\* Corresponding author. Tel: +41-21-6934510; fax: +41-21-6933604.

*E-mail address:* valerio.scarani@physics.unige.ch (V. Scarani).

PF is achieved in an EPR experiment, the EPR correlations should disappear.

In this paper we throw some light on this question by analysing long-distance EPR experiments performed between two telecom stations (Bellevue and Bernex) separated by 10.6 km [14]. The main idea is that, having observed standard EPR correlations, we are able to set a lower bound for the speed of quantum information in any given frame. The structure of the paper is as follows. In Section 2, we define the speed of quantum information, and give its transformation law under a Lorentz boost. In Section 3, we introduce a good candidate for the PF, namely, the frame of the cosmic microwave background radiation (CMB), and give its speed with respect to the rest frame of our laboratory (G-frame, where G stands for Geneva). The results of these two sections are combined in Section 4 with experimental data, leading to the announced lower bound for the speed of quantum information in the CMB-frame. Section 5 is a conclusion.

## 2. The speed of quantum information

In an optical EPR experiment (Fig. 1), two photons are produced in an entangled state and sent to two analyzing stations  $A$  and  $B$ . The quantum entanglement manifests itself by the interference fringes that are observed in the coincidence counts of the detectors in  $A$  and  $B$ . These interferences are predicted by QM; still, many physicists are not at ease with correlations that arise between two space-like separated events. The correlation are sometimes considered as due to a “superluminal influence” that the first particle to reach its detector sends to the second one. In this work, we call “speed of quantum information”  $\vec{v}_{\text{QI}}$  the superluminal speed at which this “influence” should propagate from one station to the other one. Of course, if two events are space-like separated, there is always a frame in which the two events are simultaneous ( $v_{\text{QI}} = \infty$ ), and a family of frames in which the ordering of the arrivals is inversed with respect to the laboratory frame. Therefore the supposed “superluminal influence” is a real physical process only in a preferred-frame theory, or in a theory in which the meaningful frames are not arbitrary [15]. However, the operational definition of  $\vec{v}_{\text{QI}}$  involves events (detections), that can be parametrized

by using the standard relativistic formalism; in other words, a speed of quantum information can be defined formally in any frame. Its definition goes as follows: since correlations were observed, the quantum information must have travelled the distance between the two detectors in the time interval between the two detections. Let us define the  $x$  axis as the axis linking the detector in  $A$  and the detector in  $B$ , oriented from  $A$  to  $B$ . Therefore, if the event “detection at  $A$ ” is parametrized in a given frame by  $(x_A, t_A)$ , and similarly for the event “detection in  $B$ ”, the speed of quantum information must have the same direction as and have a higher value than

$$v_{\text{QI},\min} = \frac{x_A - x_B}{t_A - t_B}. \quad (1)$$

Let  $\vec{v}$  be the speed of a given frame with respect to the laboratory frame,  $v_x$  its projection on the  $x$  axis. The Lorentz transform allows us to express  $v_{\text{QI},\min}$  as a function of the values measured in the laboratory frame:

$$\begin{cases} x = \gamma(x_{\text{Lab}} - v_x t_{\text{Lab}}), \\ t = \gamma(t_{\text{Lab}} - v_x x_{\text{Lab}}/c^2) \end{cases} \rightarrow v_{\text{QI},\min}(v_x) = -\frac{d_{AB} + v_x \tau}{\tau + v_x d_{AB}/c^2}, \quad (2)$$

where  $\gamma = (1 - v^2/c^2)^{-1/2}$ ,  $d_{AB} = x_{B,\text{Lab}} - x_{A,\text{Lab}}$ , which is positive by our convention, and  $\tau = t_{A,\text{Lab}} - t_{B,\text{Lab}}$ , which is positive if the detection at  $A$  occurs, in the laboratory frame, after the detection in  $B$ . Introducing the dimensionless parameters  $\beta_x = v_x/c$ , describing the boost, and  $r = c\tau/d_{AB} = c/|v_{\text{QI},\min}(\text{Lab})|$ , quantifying the alignment in the laboratory frame, we find the expression

$$v_{\text{QI},\min}(\beta) = -c \frac{1 + r\beta_x}{r + \beta_x}. \quad (3)$$

Thus  $v_{\text{QI},\min}(\beta)$  depends on the orientation of the setup through  $\beta_x$ , and on the *relative* precision of the alignment  $r$ . It can be checked that  $|v_{\text{QI},\min}(\beta)| > c$  if and only if  $|v_{\text{QI},\min}(\text{Lab})| > c$  ( $r < 1$ ), independent on  $\beta$ ; this is indeed a necessary property of the speed of quantum information, since two events that are space-like separated cannot be found to be time-like separated in any frame. From now on, we suppose that we consider space-like separated events ( $r < 1$ ).

Due to finite accuracy in the experiments, one can never conclude to perfect simultaneity. In an EPR

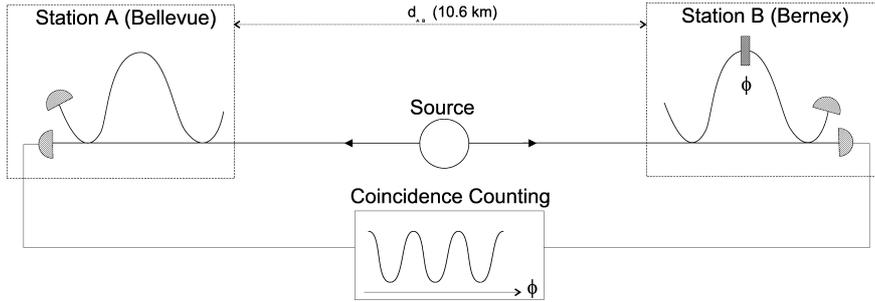


Fig. 1. Schematic representation of the experiment consisting on a photon pair source and two analyzers separated by more than 10 km.

experiment, the localization  $\Delta\tau$  of the photons is a limit to the accuracy; the best localization that can be achieved is the coherence length  $\tau_c$ . The upper limit of  $|v_{\text{QL},\min}(\beta)|$  that one can reach is thus  $d_{AB}/\tau_c$  in all frames.

So far, the discussion is valid for any inertial frame. The next section introduces a particular moving frame, the CMB-frame.

### 3. Linking the G-frame to the CMB-frame

As emphasized explicitly in several works, even recently [4,9,10], the frame of the cosmic microwave background radiation (CMB-frame) is a natural candidate for playing the role of PF. The CMB-frame is defined as the frame in which the cosmic background radiation is isotropic [16].

To find the relative speed of the G-frame with respect to the CMB-frame, we need to take into account: (i) The speed  $\vec{v}_{\text{S}|\text{CMB}}$  of the barycenter of the solar system (identified with the Sun for our purposes) with respect to the CMB; (ii) the speed  $\vec{v}_{\text{E}|\text{S}}$  of the Earth with respect to the Sun, and (iii) the spin of the Earth, giving the speed  $\vec{v}_{\text{G}|\text{E}}$  of Geneva with respect to the center of mass of the Earth. The speed  $\vec{v}_{\text{S}|\text{CMB}}$  is given in the literature [17]: its magnitude is  $v_{\text{S}|\text{CMB}} = 369$  km/s; its direction in the orthogonal celestial coordinates (see Dig. 2) is  $(\alpha = 11.20h, \delta = -7.22^\circ)$ . For the other two speeds, we have in magnitude:  $v_{\text{E}|\text{S}} = 2\pi D_{\oplus}/(1 \text{ year}) \approx 30$  km/s, where  $D_{\oplus}$  is the distance Earth–Sun, and  $v_{\text{G}|\text{E}} = 2\pi R_{\oplus} \cos l_G/(24 \text{ h}) \approx 0.3$  km/s, where  $R_{\oplus}$  is the Earth’s radius and  $l_G \approx 43^\circ$  is the latitude of

Geneva. The three speeds being much smaller than  $c$ , we can use the Galilean addition rule:

$$\vec{v}_{\text{G}|\text{CMB}} = \vec{v}_{\text{G}|\text{E}} + \vec{v}_{\text{E}|\text{S}} + \vec{v}_{\text{S}|\text{CMB}}. \quad (4)$$

Due to its magnitude, the contribution of  $\vec{v}_{\text{G}|\text{E}}$  can be neglected. Of course, the speed of the CMB-frame with respect to the G-frame is given by  $\vec{v}_{\text{CMB}|\text{G}} = -\vec{v}_{\text{G}|\text{CMB}}$ .

To perform the calculation, we need to introduce a system of coordinates, which can be chosen arbitrarily. We choose a Cartesian coordinate system defined as follows: the  $z$ -axis is the North–South axis of the Earth, oriented in the N direction. The celestial equatorial plane is therefore the  $(x, y)$  plane. In this plane, the  $x$  direction is chosen to be the direction of the vernal point, that is the point where the ecliptic intersects the celestial equatorial plane at the Spring equinox (see Fig. 2). Thus

$$\vec{v}_{\text{S}|\text{CMB}} = v_{\text{S}|\text{CMB}} \begin{pmatrix} \cos \phi_v \sin \theta_v \\ \sin \phi_v \sin \theta_v \\ \cos \theta_v \end{pmatrix}, \quad (5)$$

with  $\phi_v = 11.20h = 168^\circ$  and  $\theta_v = 97.22^\circ$  [17]. Also, neglecting the excentricity of the Earth’s orbit, one finds

$$\begin{aligned} \vec{v}_{\text{E}|\text{S}} &= -\vec{v}_{\text{S}|\text{E}} \\ &= \omega_y D_{\oplus} \begin{pmatrix} -\sin(\omega_y t + \theta_0) \\ -\cos(\omega_y t + \theta_0) \cos \theta_e \\ -\cos(\omega_y t + \theta_0) \sin \theta_e \end{pmatrix}, \end{aligned} \quad (6)$$

where  $\omega_y = 2\pi/(1 \text{ year})$  and  $\theta_e = 23.5^\circ$  is the inclination of the ecliptic plane with respect to the equatorial plane.

We still have to define the origin of time. It is natural to set  $t = 0$  at the beginning of the EPR experiment we

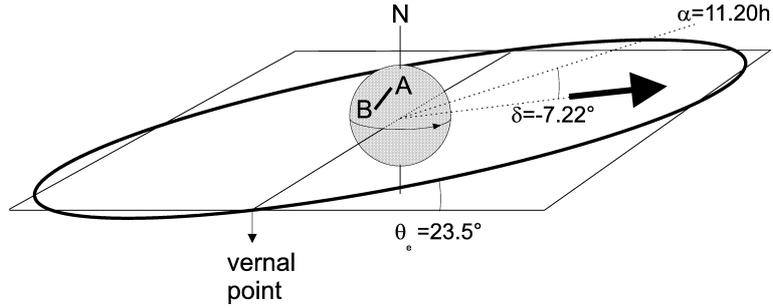


Fig. 2. Schematic representation of the direction of  $\vec{v}_{\text{S}}\text{CMB}$  (black arrow). The sphere is the Earth, A is Bellevue and B is Bernex. The plane (celestial equator) contains the equator of the Earth; the black curve is the trajectory of the Sun as seen from the Earth (ecliptic). The vernal point is defined as the intersection of the celestial equator and the ecliptic at the Spring equinox.

want to analyse. Since the acquisition time is typically some hours,  $\omega_y t$  in (6) can be set to 0. The definition of  $t = 0$  provides also the interpretation of  $\theta_0$ : this angle is defined as  $\omega_y \Delta T$ , with  $\Delta T$  the time elapsed since the Spring equinox at the moment of the experiment.

Recall that in Eq. (3) we must enter the projection of  $\vec{v}_{\text{CMB}}\text{G}$  on the direction defined by the straight line joining the two detectors. In our coordinate system, this direction is given by

$$\hat{e}_x \equiv \frac{\vec{AB}}{|\vec{AB}|} = \frac{1}{N} \begin{pmatrix} \sin \theta_B \cos \phi_B - \sin \theta_A \cos \phi_A \\ \sin \theta_B \sin \phi_B - \sin \theta_A \sin \phi_A \\ \cos \theta_B - \cos \theta_A \end{pmatrix}, \quad (7)$$

where  $N = \sqrt{2[1 - \cos \theta_A \cos \theta_B - \sin \theta_A \sin \theta_B \cdot \cos(\phi_A - \phi_B)]^{1/2}}$ . If A is Bellevue ( $46^\circ 15' \text{N}$ ,  $6^\circ 09' \text{E}$ ) and B is Bernex ( $46^\circ 10' \text{N}$ ,  $6^\circ 05' \text{E}$ ), then  $\theta_A = 43^\circ 45'$ ,  $\theta_B = 43^\circ 50'$ ,  $\phi_A = \phi_0 + \omega_d t$ ,  $\phi_B = \phi_0 + \omega_d t - 0^\circ 04'$ . The angle  $\phi_0$  measures the position of the vernal point with respect to the meridian of Bellevue at the beginning of the experiment.

#### 4. Analysis of experimental data

For the study of the speed of quantum information with an EPR setup, the precision of the alignment  $|r| = c\tau/d_{AB}$  is the figure of merit. In fact, looking at Eq. (3), we see that the simultaneity condition  $r + \beta_x = 0$  can be satisfied only if the precision of the alignment  $r$  satisfies  $|r| < \max_t |\beta_x(t)| \approx |\vec{v}|/c$ . Thus,

the smaller the speed of the considered frame with respect to the laboratory frame, the higher the precision required to satisfy the simultaneity condition. In other words, for a given frame, two situations may arise: (i) The situation of *bad alignment* is described by  $|r| > \max |\beta_x|$ . In this case,  $|v_{\text{QI},\min}(\beta)| \approx c/|r|$ . (ii) The situation of *good alignment* is the opposite one: the simultaneity condition can be satisfied.  $|v_{\text{QI},\min}(\text{CMB})|$  is no more limited by  $r$ , but there are still two possible limiting factors. The first one is the localization of the photons, that we have already mentioned. The second one is the rotation of the Earth around its axis. Due to this rotation,  $\beta_x$  varies with time, since the line Bellevue–Bernex is not parallel to this axis — and in fact,  $\hat{e}_x$  given by Eq. (7) varies with  $t$  through  $\phi_A$  and  $\phi_B$ . In particular,  $\beta_x$  is not constant during the time needed to record an interference fringe. The speed  $v_{\text{QI},\min}(\beta)$  depends on  $\beta_x$  and on  $r$ , so in principle one could keep it constant even though  $\beta_x$  changes, “simply” by performing the suitable continuous correction of the alignment  $r$ . In practice, such a protocol is cumbersome. So  $v_{\text{QI},\min}(\beta)$  also varies during the recording of a fringe. This second limiting factor is actually the most important one, as will be shortly shown.

If the considered frame is the CMB-frame, whose motion is rather slow since  $|\vec{v}_{\text{CMB}}\text{G}| \approx 300 \text{ km/s}$ , then a good alignment is obtained for  $|r| \sim 10^{-3}$ . Since such a precision was not looked for, the alignment was probably “bad” in most of the previously reported EPR experiments [18]. In the experiment that we are going to consider [14], the distance between the detectors was 10.6 km, and the difference in the two arms was

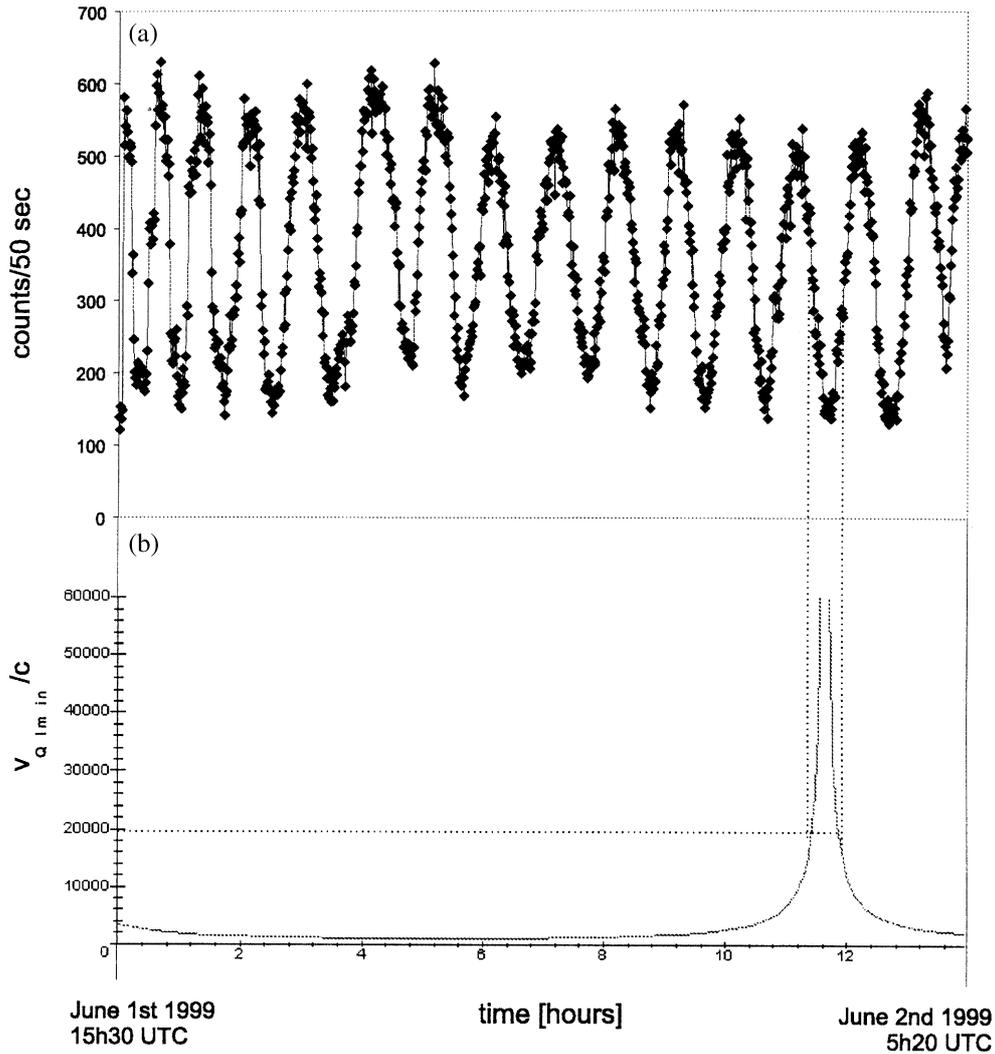


Fig. 3. (a) Two-photon interference fringes measured over 15 h, each data point corresponds to a time interval of 50 s. (b) The value of  $|v_{QI,min}(CMB)|$  calculated by Eq. (3) for the day and hours of the experiment.

lowered below 10 mm, whence  $|r| \geq 10^{-6}$ : the “good alignment” criterion for the CMB-frame is clearly fulfilled. Due to chromatic dispersion in the fibers, the localization of the photons was  $\Delta\tau = 90$  ps, so that the maximal value of  $|v_{QI,min}(CMB)|$  that we can hope to obtain is about  $3.5 \times 10^5 c$  [19]. It took about one hour to record a fringe, the detection rates being lowered by the photon losses in the fibers.

A typical EPR correlation trace is given in Fig. 3(a). This data acquisition started on June 1, 1999 at

15 h 30 min UTC (whence  $\theta_0 = 1.24$  rad and  $\phi_0 = 2.247$  rad), and ended on June 2 at 6 h 30 min UTC. Slight variations of the temperature led to a variation of the length of the fibers linking the source to the analyzing stations, that is, to a variation of the alignment  $\tau$  in the G-frame. We did not monitor  $\tau(t)$  continuously; we assume a linear interpolation between the initial value  $c\tau_i = 2$  mm, and the final value  $c\tau_f = 14$  mm. This assumption completes the set of numerical values needed to evaluate  $|v_{QI,min}(CMB)|$  according to

formula (3), as a function of time, that is, as a function the Earth’s rotation around its axis. The numerical evaluation is shown in Fig. 3(b). We see that at a given moment (about 3 h UTC) the detection events were simultaneous in the CMB-frame. Even at that moment, the visibility of the fringes is not reduced. By requiring the reduction of a whole half-fringe as a conservative criterion for fixing a limit to the speed of quantum information (see dotted lines in Fig. 3), we find the lower bound  $|v_{\text{QI,min}}(\text{CMB})| = 1.5 \times 10^4 c$ .

## 5. Conclusions

We have presented the first analysis of the results of an EPR experiment in the frame of the cosmic microwave background radiation. The conservative bound that we obtained for the “speed of quantum information” in that frame,  $|v_{\text{QI,min}}(\text{CMB})| = 2 \times 10^4 c$ , is still quite impressive, but, like most physicists, the present authors will not be astonished if further experiments provide an even higher value.

In the experiment that we analyzed, the recording of the fringes is slow compared to the variation of  $v_x$  induced by the rotation of the Earth, and this is the constraint that fixes the above limit of  $|v_{\text{QI,min}}(\text{CMB})|$ . The precision required for some planned experiments with laboratory distances [20] should increase the bound  $|v_{\text{QI,min}}(\text{CMB})|$  up to  $5 \times 10^5 c$ , the rotation of the Earth still being the most important limiting factor. As an order of magnitude, we estimated that a fringe should be completed in less than 5 s in order to reach the limit imposed by the delocalization of the photons.

The method of analysis that has been developed in this work could of course be applied to all possible frames. Under this viewpoint, the best arrangement would consist in orienting the detecting stations *A* and *B* along the *W–E* direction. In this case, due to the rotation of the Earth, in a twelve-hour scan the condition  $\beta_x = 0$  is satisfied for all frames *F* for which  $\vec{v}_{E|F}$  is constant during the scan.

## Acknowledgements

We acknowledge discussion with R. Durrer, Ch. Lineweaver, I. Percival and Ph. Eberhard. We ac-

knowledge financial support of the “Fondation Odier de psycho-physique” and the Swiss National Science Foundation.

## References

- [1] Y. Aharonov, D.Z. Albert, Phys. Rev. D 24 (1981) 359; Y. Aharonov, D.Z. Albert, Phys. Rev. D 29 (1984) 228.
- [2] H.-P. Breuer, F. Petruccione (Eds.), *Open Systems and Measurement in Relativistic Quantum Theory*, Springer, Berlin, 1999.
- [3] F. Reuse, Ann. Phys. 154 (1984) 161.
- [4] P. Caban, J. Rembielinski, Phys. Rev. A 59 (1999) 4187.
- [5] H. Bacry, *Localizability and Space in Quantum Physics*, Lecture Notes in Physics, Vol. 308, Springer, Berlin, 1988.
- [6] A. Peres, Phys. Rev. A 61 (2000) 022117.
- [7] R. Omnès, Are there unsolved problems in the interpretation of quantum mechanics?, in: H.-P. Breuer, F. Petruccione (Eds.), *Open Systems and Measurement in Relativistic Quantum Theory*, Springer, Berlin, 1999.
- [8] A. Peres, *Quantum Theory: Concepts and Methods*, Kluwer, Dordrecht, 1998.
- [9] T. Maudlin, Space-time in the quantum world, in: J.T. Cushing, A. Fine, S. Goldstein (Eds.), *Bohmian Mechanics and Quantum Theory: An Appraisal*, Kluwer, Dordrecht, 1996.
- [10] L. Hardy, Phys. Rev. Lett. 68 (1992) 2981.
- [11] Obviously, we do not discuss here historical versions of the PF (the “ether” of the XIX century), for which experiments have been designed and performed. The requirement, that a PF for quantum phenomena must not predict wrong results for classical experiments, can be fulfilled [3,4].
- [12] See Section 2 for a formal definition. We call “quantum information” what is probably better known as “spooky-action-at-a-distance”, in Einstein’s terms; or as “superluminal influences”. It must of course be kept in mind that this “information” cannot be controlled, so that in particular it cannot be used for faster-than-light communication.
- [13] Ph. Eberhard, A Realistic Model for Quantum Theory with a Locality Property, in: W. Schommers (Ed.), *Quantum Theory and Pictures of Reality*, Springer, Berlin, 1989.
- [14] H. Zbinden et al., quant-ph/0002031; H. Zbinden et al., quant-ph/0007009, submitted. Most of the measurements described in these references were performed by replacing a detector with an absorbing tape. The experiment that we analyse in this work has been performed with standard detectors in each station.
- [15] A. Suarez, V. Scarani, Phys. Lett. A 232 (1997) 9.
- [16] All the basic definitions and an up-to-date review of the experimental data about the CMB can be found online at <http://pdg.lbl.gov>.
- [17] C. Lineweaver et al., Astrophys. J. 470 (1996) 38.
- [18] As expected, a discussion of the precision of the alignment is missing in most works, because this was not a primary concern in EPR experiments. The Innsbruck group reported a precision of 1/500 for their experiment on 500 m, G. Weihs et al., Phys. Rev. Lett. 81 (1998) 5039.

For experiments in a laboratory like Aspect's "good alignment" would mean that the two arms of the interferometer should differ by at most 10 mm, A. Aspect et al., *Phys. Rev. Lett.* 49 (1982) 91.

[19] In some experiments described in [14], a filter was used to reduce the effect of the chromatic dispersion, leading to

$\Delta\tau = 5$  ps. We thus obtained in the G-frame  $|v_{\text{QI},\text{min}}(\text{Lab})| \approx \frac{2}{3} \times 10^7 c$ . For our present analysis we do not need such a precision, since the spin of the Earth is the most important limiting factor.

[20] A. Suarez, *Phys. Lett. A* 269 (2000) 293.